

KAM Theory and Applications in Celestial Mechanics – Fourth Lecture: Lower Dimensional Invariant Tori and KAM Global Point of View on the Dynamics

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The Hamiltonian of a planar “SJSU-like” system

The Hamiltonian of a planetary 4–body system writes as

$$F(\mathbf{r}, \tilde{\mathbf{r}}) = T^{(0)}(\tilde{\mathbf{r}}) + U^{(0)}(\mathbf{r}) + T^{(1)}(\tilde{\mathbf{r}}) + U^{(1)}(\mathbf{r}),$$

where \mathbf{r} are the heliocentric coordinates, $\tilde{\mathbf{r}}$ the conjugated momenta and

$$T^{(0)}(\tilde{\mathbf{r}}) = \frac{1}{2} \sum_{j=1}^3 \|\tilde{\mathbf{r}}_j\|^2 \left(\frac{1}{m_0} + \frac{1}{m_j} \right),$$

$$U^{(0)}(\mathbf{r}) = -\mathcal{G} \sum_{j=1}^3 \frac{m_0 m_j}{\|\mathbf{r}_j\|},$$

$$T^{(1)}(\tilde{\mathbf{r}}) = \frac{\tilde{\mathbf{r}}_1 \cdot \tilde{\mathbf{r}}_2}{m_0} + \frac{\tilde{\mathbf{r}}_1 \cdot \tilde{\mathbf{r}}_3}{m_0} + \frac{\tilde{\mathbf{r}}_2 \cdot \tilde{\mathbf{r}}_3}{m_0},$$

$$U^{(1)}(\mathbf{r}) = -\mathcal{G} \left(\frac{m_1 m_2}{\|\mathbf{r}_1 - \mathbf{r}_2\|} + \frac{m_1 m_3}{\|\mathbf{r}_1 - \mathbf{r}_3\|} + \frac{m_2 m_3}{\|\mathbf{r}_2 - \mathbf{r}_3\|} \right).$$

We fixed the masses and the average semi-major axes so to be equal to those of the real Sun–Jupiter–Saturn–Uranus system.

The Poincaré variables in the plane

$$\Lambda_j = \frac{m_0 m_j}{m_0 + m_j} \sqrt{\mathcal{G}(m_0 + m_j) a_j} \quad \lambda_j = M_j + \omega_j$$

$$\xi_j = \sqrt{2\Lambda_j} \sqrt{1 - \sqrt{1 - e_j^2}} \cos(\omega_j) \quad \eta_j = -\sqrt{2\Lambda_j} \sqrt{1 - \sqrt{1 - e_j^2}} \sin(\omega_j)$$

where a_j , e_j , M_j and ω_j are the orbital elements of the j -th planet.

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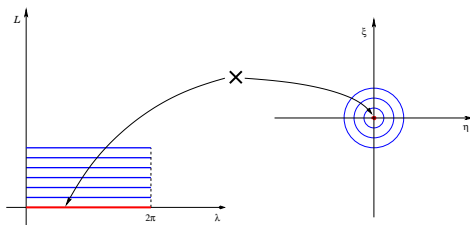
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Let us introduce new actions $\mathbf{L} = \mathbf{\Lambda} - \mathbf{\Lambda}^*$, where Λ_j^* is calculated with the average value a_j^* of the semi-major axis of the j -th planet.

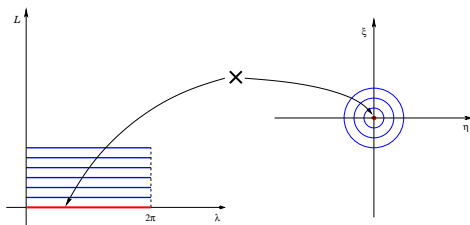
Elliptic tori as “extensions” of equilibrium points

- **Question:** what is *the equilibrium point of the secular part* (that is a 3 d.o.f. system) *with respect to the flow of the averaged Hamiltonian* $\langle F \rangle_\lambda$ (6 d.o.f. system)?
- **Answer:** **an elliptic torus!**



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- **Problem:** **can we locate elliptic tori in the complete (i.e. non-averaged) Hamiltonian F ?**

Elliptic tori in (already seen) “resonant regions”

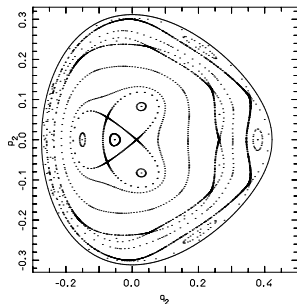


Figure: Poincaré sections for the Hénon–Heiles model

$H(\mathbf{p}, \mathbf{q}) = \omega_1(p_1^2 + q_1^2)/2 + \omega_2(p_2^2 + q_2^2)/2 + q_1^2 q_2 - q_2^3/3$ with $\omega_1 = 1$ and $\omega_2 = (\sqrt{5} - 1)/2$. The energy level is fixed so that $E = 0.030$. In this case the escape energy value is $E_e = 0.03934466$.

- **Remark:** in Poincaré sections, elliptic tori, that are invariant with respect to the flow of H , are seen as single points contoured by closed 1D–curves. *In figure above, they are visible in the so called “chains of ordered islands”.*

The wanted normal form

Let us imagine to have already performed an infinite sequence of can. transf. so to bring the Hamiltonian in the wanted normal form

$$H^{(\infty)}(\mathbf{L}, \boldsymbol{\lambda}, \boldsymbol{\xi}, \boldsymbol{\eta}) = \boldsymbol{\omega}^{(\infty)} \cdot \mathbf{L} + \sum_{j=1}^{n_2} \frac{\Omega_j^{(\infty)}}{2} (\xi_j^2 + \eta_j^2) + \mathcal{R}^{(\infty)}(\mathbf{L}, \boldsymbol{\lambda}, \boldsymbol{\xi}, \boldsymbol{\eta}),$$

where the frequency vectors $\boldsymbol{\omega}^{(\infty)} \in \mathbb{R}^{n_1}$ and $\boldsymbol{\Omega}^{(\infty)} \in \mathbb{R}^{n_2}$, with $n_1 + n_2 = n$, being n the number of d.o.f. (in our model $n_1 = n_2 = 3$). Moreover, the remainder term is such that

$$\mathcal{R}^{(\infty)}(\mathbf{L}, \boldsymbol{\lambda}, \boldsymbol{\xi}, \boldsymbol{\eta}) = \mathcal{O}(\|\mathbf{L}\|^2) + \mathcal{O}(\|\mathbf{L}\| \|\boldsymbol{\xi}, \boldsymbol{\eta}\|) + \mathcal{O}(\|\boldsymbol{\xi}, \boldsymbol{\eta}\|^3).$$

$\forall (\mathbf{0}, \boldsymbol{\lambda}, \mathbf{0}, \mathbf{0}) \in (\mathbf{0}, \mathbb{T}^{n_1}, \mathbf{0}, \mathbf{0})$, the Hamilton eq.s can be easily solved

$$\dot{\mathbf{L}} = \mathbf{0}, \quad \dot{\boldsymbol{\lambda}} = \boldsymbol{\omega}^{(\infty)}, \quad \dot{\boldsymbol{\xi}} = \mathbf{0}, \quad \dot{\boldsymbol{\eta}} = \mathbf{0}.$$

Thus, **the flow induced by $H^{(\infty)}$ on the invariant lower-dimensional torus $(\mathbf{0}, \mathbb{T}^{n_1}, \mathbf{0}, \mathbf{0})$ is quasi-periodic with frequency vector $\boldsymbol{\omega}^{(\infty)}$, while $\boldsymbol{\Omega}^{(\infty)}$ is the limit frequency vector of the small oscillations transversal to the elliptic torus.**

The normal form Hamiltonian up to order $r - 1$

Where it is convenient, we refer to the secular variables with action–angle coordinates (\mathbf{I}, φ) such that $\xi_j = \sqrt{2I_j} \cos \varphi_j$, $\eta_j = \sqrt{2I_j} \sin \varphi_j$, $\forall j = 1, \dots, n_2$. The Hamiltonian F representing our planar “SJSU-like” model can be written in the following general form (with $r = 1$):

$$\begin{aligned}
 H^{(r-1)} = & \boldsymbol{\omega}^{(r-1)} \cdot \mathbf{L} + \boldsymbol{\Omega}^{(r-1)} \cdot \mathbf{I} + \sum_{j_1 \geq 2} h_{j_1,0}(\mathbf{L}) + \sum_{s \geq r} f_{0,0}^{(r-1,s)}(\boldsymbol{\lambda}) + \\
 & \sum_{s \geq r} f_{0,1}^{(r-1,s)}(\boldsymbol{\lambda}, \boldsymbol{\xi}, \boldsymbol{\eta}) + \sum_{s \geq r} f_{1,0}^{(r-1,s)}(\mathbf{L}, \boldsymbol{\lambda}) + \sum_{s \geq r} f_{0,2}^{(r-1,s)}(\boldsymbol{\lambda}, \boldsymbol{\xi}, \boldsymbol{\eta}) \\
 & + \sum_{2j_1+j_2 \geq 3} \sum_{s > 0} f_{j_1,j_2}^{(r-1,s)}(\mathbf{L}, \boldsymbol{\lambda}, \boldsymbol{\xi}, \boldsymbol{\eta}),
 \end{aligned}$$

where $r - 1$ means the normalization step, $h_{j_1,0}$ are homogeneous polynomials of degree j_1 in \mathbf{L} , $f_{j_1,j_2}^{(r-1,s)}$ are hom. pol. of degree j_1 and j_2 in \mathbf{L} and in $(\boldsymbol{\xi}, \boldsymbol{\eta})$, resp., while they are also trig. pol. of degree at most $2s$ in the angles $\boldsymbol{\lambda}$. Thus, *in the previous formula, each term has a finite Taylor–Fourier expansion. The normalization procedure has to eliminate the (red color) perturbing terms.*

Normalization procedure: the homological equations

$$\left\{ \chi_0^{(r)}, \omega^{(r-1)} \cdot \mathbf{L} \right\} + \sum_{s=1}^r f_{0,0}^{(r-1,s)}(\boldsymbol{\lambda}) = 0 ,$$

$$\left\{ \chi_1^{(r)}, \omega^{(r-1)} \cdot \mathbf{L} + \boldsymbol{\Omega}^{(r-1)} \cdot \mathbf{I} \right\} + \sum_{s=0}^r f_{0,1}^{(I;r,s)}(\boldsymbol{\lambda}, \boldsymbol{\xi}, \boldsymbol{\eta}) = 0 ,$$

$$\left\{ \chi_2^{(r)}, \omega^{(r-1)} \cdot \mathbf{L} \right\} + \sum_{s=1}^r f_{1,0}^{(II;r,s)}(\mathbf{L}, \boldsymbol{\lambda}) = 0 ,$$

$$\left\{ Y_2^{(r)}, \omega^{(r-1)} \cdot \mathbf{L} + \boldsymbol{\Omega}^{(r-1)} \cdot \mathbf{I} \right\} + \sum_{s=1}^r f_{0,2}^{(II;r,s)}(\boldsymbol{\lambda}, \boldsymbol{\xi}, \boldsymbol{\eta}) = 0 ,$$

$$\left\{ D_2^{(r)}, \boldsymbol{\Omega}^{(r-1)} \cdot \mathbf{I} \right\} + f_{0,2}^{(II;r,0)}(\mathbf{I}, \boldsymbol{\varphi}) - \langle f_{0,2}^{(II;r,0)} \rangle_{\boldsymbol{\varphi}} = 0 .$$

Normalization procedure for elliptic tori: main ideas

- Each r -th normalization step is performed by *composing of three canonical transformations* $\exp \mathcal{L}_{\chi_0^{(r)}}$, $\exp \mathcal{L}_{\chi_1^{(r)}}$ and $\exp \mathcal{L}_{\chi_2^{(r)}}$, where *the generating functions* $\chi_0^{(r)}$, $\chi_1^{(r)}$ and $\chi_2^{(r)}$ **are determined so to eliminate the perturbing terms**
 - *independent from both \mathbf{L} and (ξ, η) ,*
 - *independent from \mathbf{L} and of degree 1 in (ξ, η) ,*
 - **either of degree 1 in \mathbf{L} and independent from (ξ, η) , or independent from \mathbf{L} and of degree 2 in (ξ, η) ,**

respectively; moreover, *each generating function “kills” perturbing terms up to trigonometric degree $2r$ in λ .*
- at the end of each normalization step, there are some terms $\mathcal{O}(\|\mathbf{L}\|)$ and $\mathcal{O}(\|(\xi, \eta)\|^2)$ that cannot be eliminated (because they do not depend on the angles); thus, they are included in the normal form terms and they induce small corrections of the frequency vectors, that are $\omega^{(r-1)} \rightarrow \omega^{(r)}$ and $\Omega^{(r-1)} \rightarrow \Omega^{(r)}$.
- *If the sequence $\{(\omega^{(r)}, \Omega^{(r)})\}_{r \geq 0}$ stays away enough from the resonances, **the normalization procedure works!** (see Sansottera, Locatelli & Giorgilli 2011).*

Testing the construction of the normal form for elliptic tori

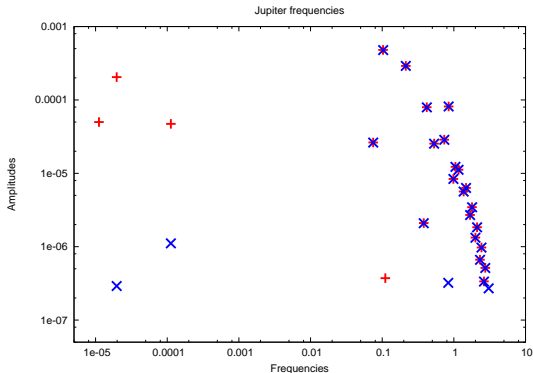


Figure: Fourier spectra with the first 30 main components of the signal

$t \rightarrow \xi_1(t) + i\eta_1(t)$ related to the *secular motion of Jupiter*, i.e.

$$\xi_1(t) + i\eta_1(t) \simeq \sum_{s=1}^{30} c_{1,s} \exp[i(\nu_{1,s}t)].$$

Frequencies $\nu_{1,s}$ and amplitudes $|c_{1,s}|$ are reported on the abscissas and the ordinates, resp. + symbols refer to a motion starting from “real initial conditions” of the “planar SJSU”, while x symbols are for an approximated elliptic torus after 9 normalization steps.

Frequency analysis shows that we are approaching an elliptic torus!

Testing the construction of the normal form for elliptic tori

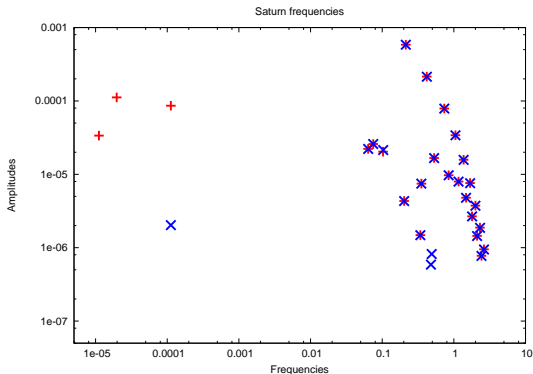


Figure: Fourier spectra with the first 30 main components of the signal

$t \rightarrow \xi_2(t) + i\eta_2(t)$ related to the *secular motion of Saturn*, i.e.

$\xi_2(t) + i\eta_2(t) \simeq \sum_{s=1}^{30} c_{2,s} \exp[i(\nu_{2,s}t)]$. Frequencies $\nu_{2,s}$ and amplitudes $|c_{2,s}|$

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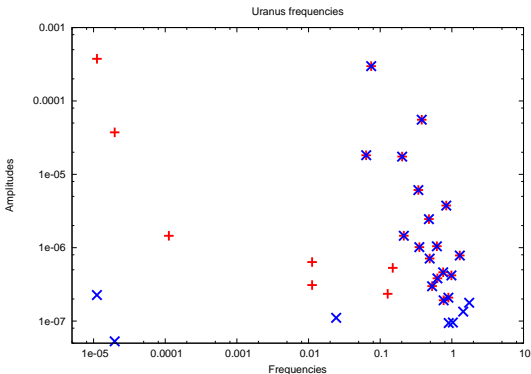


Figure: Fourier spectra with the first 30 main components of the signal

$t \rightarrow \xi_3(t) + i\eta_3(t)$ related to the *secular motion of Uranus*, i.e.

$\xi_3(t) + i\eta_3(t) \simeq \sum_{s=1}^{30} c_{3,s} \exp[i(\nu_{3,s}t)]$. Frequencies $\nu_{3,s}$ and amplitudes $|c_{3,s}|$

are reported on the abscissas and the ordinates, resp. + symbols refer to a motion starting from “real initial conditions” of the “planar SJSU”, while x symbols are for an approximated elliptic torus after 9 normalization steps.

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Main differences in the construction of the Kolmogorov's normal form: KAM tori vs. elliptic lower dimensional tori

- The small parameter is the mass ratio $\max_{j=1,2,3} m_j / m_0$.
- At each normalization step for elliptic tori, we need a non-resonance condition of type

$$\min_{0 < |\mathbf{k}| \leq 2r} \left| \mathbf{k} \cdot \boldsymbol{\omega}^{(r-1)} \right| > 0 ,$$

that is similar to that needed by the usual KAM tori, but we need also the so called **Melnikov's conditions**:

$$\min_{|\mathbf{k}| \leq 2r} \min_{1 \leq |\ell| \leq 2} \left| \mathbf{k} \cdot \boldsymbol{\omega}^{(r-1)} + \ell \cdot \boldsymbol{\Omega}^{(r-1)} \right| > 0 .$$

- In case of elliptic tori, **the frequencies are not fixed "a priori"**, and the algorithm let them (very slightly) change at each step. Thus, eventual resonances can show up after some normalization step.
- The system is usually parameterized with respect to n_1 independent variables (e.g., the average semi-major axes where the initial expansions are centered about). Thus, ***the final result must hold true on a set having positive Lebesgue measure. This is really in the spirit of the original Arnold's proof scheme.***

Arnold's proof scheme of KAM theorem



At the first normalization step, you cut out a first group of main resonances from the phase space and you perform a canonical transformation so to eliminate those resonant perturbing terms from the Hamiltonian, which is lead to the form:

$$H^{(1)}(\mathbf{p}, \mathbf{q}) = h^{(1)}(\mathbf{p}) + \mathcal{R}^{(1)}(\mathbf{p}, \mathbf{q}) ,$$

where $\mathcal{R}^{(1)}$ is a small remainder term.

Arnold's proof scheme of KAM theorem



At the second normalization step, you cut out a second group of main resonances from the phase space and you perform another canonical transformation so to write the Hamiltonian in the form:

$$H^{(2)}(\mathbf{p}, \mathbf{q}) = h^{(2)}(\mathbf{p}) + \mathcal{R}^{(2)}(\mathbf{p}, \mathbf{q}) ,$$

where $\|\mathcal{R}^{(2)}\| = \mathcal{O}(\|\mathcal{R}^{(1)}\|^2)$ is a remainder term much smaller than the previous one. *After having iterated infinitely many times this procedure the Hamiltonian is convergent (if the initial perturbation is small enough) on a Cantor set of invariant tori with positive Lebesgue measure.*

Arnold web in numerics for a 4D-symplectic map

The *Coupled Rational Shifted Standard Map (CRSSM)* is defined so that

$$y_1' = y_1 + \varepsilon_1 f_1(x_1) + \gamma_+ f_3(x_1 + x_2) + \gamma_- f_3(x_1 - x_2),$$

$$y_2' = y_2 + \varepsilon_2 f_2(x_2) + \gamma_+ f_3(x_1 + x_2) - \gamma_- f_3(x_1 - x_2),$$

$$x_1' = x_1 + \varepsilon_1 y_1' \quad \text{mod } 2\pi ,$$

$$x_2' = x_2 + \varepsilon_2 y_2' \quad \text{mod } 2\pi ,$$

where angles $x_i \in [0, 2\pi)$, actions $y_i \in [0, 2\pi/\varepsilon_i)$, $\forall i = 1, 2$, while the *perturbing functions are such that $f_i(x) = -\sin x / (1 - \mu_i \cos x)$* ,

$\forall i = 1, 2, 3$, being $\varepsilon_1, \varepsilon_2, \gamma_{\pm}, \mu_1, \mu_2$ and μ_3 *fixed small parameters*.

- Consider a regular grid of initial values $(y_{1;0}, y_{2;0})$ of the actions;
- for each initial condition of type $(y_{1;0}, y_{2;0}, 0, 0)$, iterate N times the map, so to produce a finite sequence of points $(y_{1;j}, y_{2;j}, x_{1;j}, x_{2;j})$ with $j = 0, \dots, N$;
- make the frequency analysis of the signals $j \rightarrow \sqrt{2y_{1;j}} \exp(ix_{1;j})$ and $j \rightarrow \sqrt{2y_{2;j}} \exp(ix_{2;j})$ by separating two "windows": $j \in [0, N/2]$ and $j \in [N/2, N]$. Draw the variations of the main frequencies as a function of the initial actions $(y_{1;0}, y_{2;0})$.

Arnold web in numerics for a 4D-symplectic map

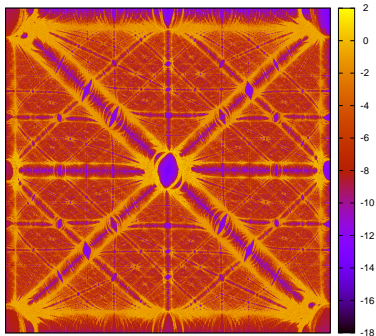


Figure: Color plot describing the Arnold web for the **CRSSM** with $\varepsilon_1 = 0.1$, $\varepsilon_2 = 0.2$, $\mu_1 = \mu_2 = \mu_3 = 0.5$, $\gamma_+ = 0.1$, $\gamma_- = 0.05$. Initial values of actions $y_{1;0}$ and $y_{2;0}$ are on abscissas and on ordinates, respectively. For each point the corresponding variation of the frequencies is reported by the color scale on the right.

Frequency analysis as a chaos indicator: *lighter colors mean chaotic motions, while darker ones are related to invariant tori.*

Diffusion along the Arnold web

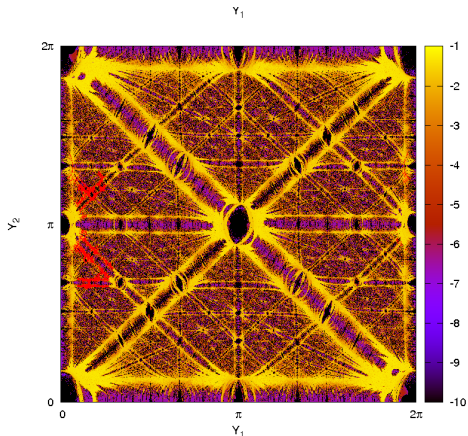


Figure: Evolution of a chaotic orbit (look at red points) along the Arnold web. *The diffusion is faster in larger resonant regions. When some resonant regions cross each other, motions can pass from one to another.*

Some more numerical experiments: the Sup-Map graph

The **standard map** is $y' = y + \varepsilon \sin x$; $x' = x + y' \pmod{2\pi}$.

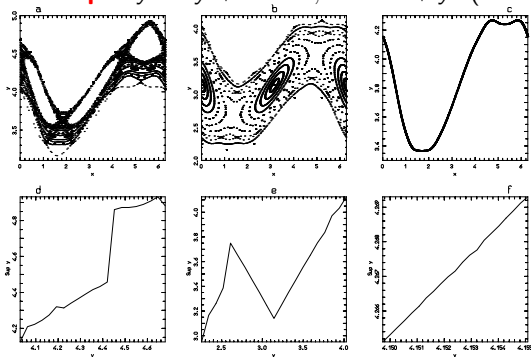


Figure: Archetypical situations occurring in a Sup-Map graph. Boxes a–c contain plots of some orbits of the standard map with $\varepsilon = 0.9$. In boxes d–f, the values of the sup of an orbit as a function of the action y_0 , being $(0, y_0)$ the initial point starting from which the orbit is generated by iterations of the standard map. The pair of **boxes (a,d), (b,e) and (c,f) refer to the vertical crossing of the chaotic zone close to an hyperbolic point, of the resonant region close to an elliptic point and to a set of invariant tori, resp.**

Some more numerical experiments: the Sup-Map graph

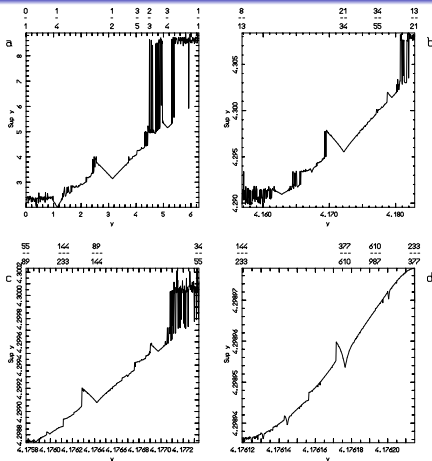


Figure: Sup-Map graph of the standard map with $\varepsilon = 0.9$. Boxes a–d are enlargements focused in a neighborhood smaller and smaller of the “golden mean” torus, that is related to the frequency $(\sqrt{5} - 1)/2$. On the top of each box, the location of the main resonances of type F_k/F_{k+1} is reported (being $\{F_k\}_{k \geq 0}$ the sequence of Fibonacci’s numbers).

Dynamics in the neighborhood of an invariant torus

Let us focus on a neighborhood of a KAM torus characterized by a Diophantine frequency vector ω , that is such that $|\mathbf{k} \cdot \omega| > \gamma/|\mathbf{k}|^\tau$ with some fixed $\gamma > 0$ and $\tau \geq n - 1$. By constructing a Birkhoff's normal form starting from the Kolmogorov's normal form and using Nekhoroshev's estimates, we can prove the following results.

- **Let $V(\varrho)$ be the measure of the complementary set with respect to the invariant tori staying at a distance smaller than ϱ from the “ ω -torus”, then**

$$V(\varrho) \simeq \exp \left(- \left(\frac{\varrho_*}{\varrho} \right)^{1/(\tau+1)} \right).$$

- **Let T_ϱ be the diffusion time needed by a motion to double its initial distance ϱ from the “ ω -torus”, then**

$$T_\varrho \simeq \exp \left(C \exp \left(\frac{1}{2n} \left(\frac{\varrho_*}{\varrho} \right)^{1/(\tau+1)} \right) \right),$$

being C and ϱ_* suitable positive constants (see *Morbidelli & Giorgilli 1995*).